Heat kernel measures on random surfaces

SEMYON KLEVTSOV AND STEVE ZELDITCH

The heat kernel on the symmetric space of positive definite Hermitian matrices is used to endow the spaces of Bergman metrics of degree k on a Riemann surface M with a family of probability measures depending on a choice of the background metric. Under a certain matrix-metric correspondence, each positive definite Hermitian matrix corresponds to a Kähler metric on M . The one and two point functions of the random metric are calculated in a variety of limits as k and t tend to infinity. In the limit when the time t goes to infinity the fluctuations of the random metric around the background metric are the same as the fluctuations of random zeros of holomorphic sections. This is due to the fact that the random zeros form the boundary of the space of Bergman metrics.

1. Introduction

In a recent series of articles [8, 9], the authors have been investigating a new approach to defining 'random surfaces'. The main idea is to define integrals over the infinite dimensional space $\mathcal{K}_{[\omega_0]}$ of metrics of fixed area 2π in a fixed conformal class $[\omega_0]$ on a Riemann surface M as limits

(1)
$$
\int_{\mathcal{K}_{[\omega_0]}} F(g) e^{-S(g)} \mathcal{D}g := \lim_{k \to \infty} \int_{\mathcal{B}_k} F_k(g) e^{-S_k(g)} \mathcal{D}_k g
$$

of integrals over finite dimensional spaces \mathcal{B}_k of Bergman metrics. Given a background metric¹ ω_0 and a choice of a basis $\{s_j(z)\}\$ of holomorphic sections of $L^k \to M$, the spaces \mathcal{B}_k can be identified with the non-positively curved symmetric space $\mathcal{P}_{N_k} := SL(N_k,\mathbb{C})/SU(N_k)$ of positive definite Hermitian matrices. The general question is to find sequences $\{d\mu_k =$ $e^{-S_k(g)}\mathcal{D}_k g$ of measures on \mathcal{B}_k which are independent of the choice of the basis $\{s_i(z)\}\$, which vary in a simple way under the change of the reference point $\omega_0 \in \mathcal{K}_{[\omega_0]}$ and have good asymptotic properties as $k \to \infty$. It would

¹With some abuse of notation we make no distinction everywhere between the metric g and the corresponding Kähler form ω , connected as $\omega = ig_{a\bar{b}}dz^a \wedge d\bar{z}^{\bar{b}}$.

be particularly interesting to construct a sequence $\{d\mu_k\}$ which tends to Liouville theory measure on metrics of fixed area $\int_M \omega_0$, although that is not the sim of the present orticle the aim of the present article.

The sequence of measures we study in this article are the heat kernel measures

(2)
$$
d\mu_k^t(P) := p_k(t, I, P)dV(P),
$$

where $dV(P)$ is Haar measure, $p_k(t, P_1, P_2)$ is the heat kernel of the symmetric space \mathcal{P}_{N_k} and I is the identity matrix. Under the matrix-metric identification $\mathcal{B}_k \simeq \mathcal{P}_{N_k}$ the identity matrix corresponds to the background metric ω_{ϕ_I} and the heat kernel measure (2) is transported to \mathcal{B}_k . The measure is invariant under the action of the unitary group $U(N_k)$. Hence it is invariant of the choice of the basis of sections used to identify metrics and matrices. Then (2) is the probability measure on \mathcal{B}_k induced by Brownian motion on \mathcal{P}_{N_k} starting at the identity I and continuing up to time t. The heat kernel measure is almost canonical, the only choices being the time t and the background metric ω_{ϕ_I} used to make the identification and to start the Brownian motion. The purpose of this article is to study the behavior of the heat kernel measure (2) on \mathcal{B}_k as $k \to \infty$. The main geometric quantities we study are the area statistics

(3)
$$
X_U(\omega) = \int_U \omega
$$

measuring the area of an open set $U \subset M$ with respect to the random area form $\omega \in \mathcal{B}_k$. We determine the means and variances of these random variables and their smooth analogues $X_f(\omega) = \int_M f \omega$ with $f \in C^{\infty}(M)$ in various regimes, e. g. when the time $t = t_k$ is allowed to vary with k. The calculations are valid for any choice of background metric and the dependence on the background metric is simple and explicit.

The heat kernel measure (2) is $U(N_k)$ -invariant in the P variable. Such invariant measures have generic form $d\mu_{\mathcal{B}_k}(P) := \mathcal{F}_{\mathcal{B}_k}(e^{\lambda}) d\mu_{\text{Haar}}(P)$, where $\mathcal{F}_{\mathcal{B}_{k}}(e^{\lambda})$ is a function of the eigenvalues of P. It was shown in [9] that the eigenvalue density $\mathcal{F}_{\mathcal{B}_k}(e^{\lambda})$ induces a function $\mathcal{F}_{k,2}(\nu_1, \nu_2)$ on \mathbb{R}^2_+ , so that the 2-point correlation function has the form,

(4)
$$
K_{2,k}(z_1, z_2) := \mathbf{E}_k \phi_P(z_1) \phi_P(z_2) = \phi_I(z_1) \phi_I(z_2) + \frac{1}{k^2} I_{2,k}(\rho),
$$

where

(5)
$$
\rho(z_1, z_2) = \frac{|B_k(z_1, z_2)|^2}{B_k(z_1, z_1)B_k(z_2, z_2)}
$$

is an important invariant of the Szegö kernel $B_k(z_1, z_2)$ of the background metric, known as the Berezin kernel. Thus, $I_{2,k}(z_1, z_2)$ is the bi-potential of the variance of the area forms (or Kähler metrics in higher dimensions) relative to the exterior tensor product $\omega_0 \boxtimes \omega_0$,

(6)
$$
\mathbf{Var}(\omega_{\phi}) = \mathbf{E}(\omega_{\phi} \boxtimes \omega_{\phi}) - \mathbf{E}(\omega_{\phi}) \boxtimes \mathbf{E}(\omega_{\phi}) = \mathbf{E}(\omega_{\phi} \boxtimes \omega_{\phi}) - \omega_0 \boxtimes \omega_0,
$$

in the sense that

(7)
$$
\mathbf{Var}(\omega_{\phi}) = \frac{1}{k^2} (i \partial \bar{\partial})_z (i \partial \bar{\partial})_w I_{2,k}(z,w), .
$$

The general formula for $I_{2,k}(\rho)$ for any $U(N_k)$ -invariant measure is,

(8)
$$
I_{2,k}(\rho) = \frac{1}{2} \int_{\mathbb{R}_+^2} \int_0^{\pi} \log(\nu_1^2 \cos^2 \beta + \nu_2^2 \sin^2 \beta)
$$

$$
\log \frac{A + \sqrt{A^2 - B^2}}{2} \mathcal{F}_{2,k}(\nu) \sin \beta \, d\beta \, d\nu_1 \, d\nu_2,
$$

with

$$
\begin{cases}\nA = (\nu_1^2 \cos^2 \beta + \nu_2^2 \sin^2 \beta)\rho + (\nu_1^2 \sin^2 \beta + \nu_2^2 \cos^2 \beta)(1 - \rho), \\
B = 2(\nu_1^2 - \nu_2^2)\sqrt{\rho(1 - \rho)} \cos \beta \sin \beta.\n\end{cases}
$$

The transform $\mathcal{F}_{\mathcal{B}_k}(e^{\lambda}) \to \mathcal{F}_{2,k}(\nu)$ is very difficult to evaluate, and we do not know how to do so directly even for the heat kernel measure. The first term of (4) is the potential of $\omega_{\phi_I}(z_1)\omega_{\phi_I}(z_2)$ where ω_{ϕ_I} is the background metric, and the second term $I_{2,k}(\rho)$ is the correction to this term, which we call the variance term. The key point is that $I_{2,k}(\rho)$ is a function only of the variable ρ . This result defnes a transform

$$
\mathcal{F}_{\mathcal{B}_k}(e^{\lambda}) \to I_{2,k}(\rho)
$$

from eigenvalue densities to variance terms depending only on ρ . It would be interesting to know if this transform is invertible in some sense, so that one can construct $U(N_k)$ -invariant measures with prescribed pair correlation functions. In this article we calculate $I_{2,k}(\rho)$ when $\mathcal{F}_{\mathcal{B}_k}(e^{\lambda})$ comes from the heat kernel measure, by a different method (also used in [9]).

1.1. Main results

In the case of heat kernel measures we calculate the pair correlation function explicitly not just for Riemann surfaces, but for general projective Kähler manifolds (see §4.2). The calculations give an explicit formula for the variance term of the 2 point function $I_{2,k}(t, \rho)$:

(9)
$$
\partial_{\rho}I_{2,k}(t,\rho) = \frac{2t}{\rho} - \frac{e^{-t/2}}{\sqrt{2\pi t}} \frac{\sqrt{1-\rho}}{\rho} \int_{-\infty}^{\infty} d\lambda \frac{e^{-\frac{1}{2t}\lambda^2} \cosh \lambda}{\sqrt{\coth^2 \lambda - \rho}}}{\log \frac{\sqrt{\coth^2 \lambda - \rho} + \sqrt{1-\rho}}{\sqrt{\coth^2 \lambda - \rho} - \sqrt{1-\rho}}}.
$$

We do not integrate the result because $I_{2,k}(t, \rho)$ is the expected "bi-potential" and the fluctuations of the metric are obtained by differentiating it. An important aspect of (9) is that the expression has no k-dependence, except for the variable ρ (5), which has the form $e^{-kD(z_1,z_2)}$ where $D(z_1,z_2)$ is the diastasis (an analog of distance-squared function for Kähler manifolds) between the points, with respect to the background metric, see $\S2$, Eq. (18).

We consider several limits of this joint formula in §4. From the geometric viewpoint, the most natural scaling of the time variable is $t_k = \varepsilon_k^{-2} t$ so that the excursion distance of the Brownian motion in \mathcal{P}_{N_k} at time t_k is essentially distance t in the Mabuchi metric on $\mathcal{K}_{[\omega_0]}$.

Similarly to $[17, 18]$, we also study the large k asymptotics both in the unscaled and scaled regimes. The scaling limit is common in related problems in the physics of N particles where one lets the number $N \to \infty$ and the volume of the surface $V \to \infty$ such that $\frac{N}{V}$ tends to a limiting density (see e.g. [14] for a similar scaling in the quantum Hall effect). The natural length scale for metrics in the Bergman space \mathcal{B}_k is $\frac{1}{\sqrt{k}}$. We consider pairs $(z_1, z_2) \in M \times M$ to be close to the diagonal if $d(z_1, z_2) \leq \frac{\log k}{\sqrt{k}}$, and to be 'off-diagonal' if $d(z_1, z_2) \geq C \frac{\log k}{\sqrt{k}}$, where $d(z_1, z_2)$ is the distance between points relative to the background metric. In the scaling limit we consider the asymptotics of $I_{2,k}(t, \rho)$ for pairs of points of the form $(z, z + \frac{u}{\sqrt{k}})$ with $|u| \leq C \log k$, in which case $\rho \simeq e^{-|u|^2}$. The scaling asymptotics combined with the time scaling $t_k \to \infty$ has a limit correlation function with a logarithmic singularity along the diagonal $z_1 = z_2$ (where $\rho = 1$). The variance $(i\partial\partial)_{z_1} (i\partial\partial)_{z_2} I_{2,k}(t,\rho)$ of the Kähler metric then has a $\delta(z_1 - z_2)$ singularity along the diagonal. In fact, the scaling limit correlation function turns out to be identical to that for zeros of random holomorphic sections determined in [17] (Lemma 3.7).

When $t \in \mathbb{R}_+$ is finite and fixed, then $I_{2,k}(t, \rho)$ is smooth at $\rho = 1$ (see (38)), and has a convergent expansion $(i\partial\bar{\partial})_{z_1} (i\partial\bar{\partial})_{z_2} I_{2,k}(t,\rho) \sim |z_1 - z_2|$ $|z_2|^2$ and there exist coefficients $a_n(t)$ so that

(10)
$$
I_{2,k}(t,\rho) = \sum_{n=0}^{\infty} a_n(t)\rho^n.
$$

Off the diagonal, $\rho \rightarrow 0$ and we can obtain the asymptotics by Taylor expanding the amplitude of the integral (9). The first term $\frac{2t}{\rho}$ is singular as $\rho \to 0$. But in §4.3, resp. in §4.4, it is shown that the $\frac{2t}{\rho}$ 'singularity' cancels in the sum of the two terms. Hence the variance of the potential is exponentially decreasing off the diagonal.

1.2. Comparison with [17, 18]

In §4.4 we first let $t \to \infty$ and then let $k \to \infty$. It turns out that in this limit, the random metrics we obtain are identical with random zero sets of holomorphic sections of the kth power L^k of the line bundle with Chern class $[\omega_0]$. As explained in §5, as $t \to \infty$ the mass of the heat kernel concentrates on the ideal boundary of the symmetric space, where the metrics correspond to the zero sets of holomorphic sections. On a Riemann surface, the random metrics become normalized sums of delta functions on random point configurations with k points. We verify that the pair correlation function of random Kähler metrics in \mathcal{B}_k in the limit $t \to \infty$ is given by the same formula as in [17] for correlations between zeros of random sections. For large t (depending on k), the random metric is close to such a point configuration measure.

We now explain this similarity in more detail so that the notation and purpose of this article are synchronized with those of [17, 18]. The model of Gaussian random holomorphic sections and the results are described in more detail in §5.1. In those articles, the role of the area form (in complex dimension one) is played by the zero set measure Z_s of a random section $s \in H^0(M, L^k)$, which defines a "singular metric". Hence it is not surprising that there are relations between random smooth metrics and random zero sets.

In [17, 18], the zero current is given by $Z_s = \frac{i}{\pi} \partial \bar{\partial} \log |s|^2$ and, analogously to (6) the variance current of zeros is defined by,

(11)
$$
\text{Var}\big(Z_{s^k}\big) = \mathbf{E}\big(Z_{s^k} \boxtimes Z_{s^k}\big) - \mathbf{E}\big(Z_{s^k}\big) \boxtimes \mathbf{E}\big(Z_{s^k}\big).
$$

In [17], it was shown that the bi-potential $I_{2,k}$ of the variance² (11) in the sense of (7) is given by dilogarithm

(12)
$$
I_{2,k}(z,w) = -\int_0^{\rho} \frac{\log(1-s)}{s} ds.
$$

In §4.4 we show that the $t \to \infty$ limit of the heat kernel ensemble gives precisely the same bi-potential, hence the same variance.

For zeros of random holomorphic sections over a Riemann surface, the area statistic (3) with respect to Z_s counts the number of zeros of s in U and is denoted by \mathcal{N}_U in [17]. it is shown there that

$$
\text{Var}(\mathcal{N}_U) = -\int_{\partial U \times \partial U} \bar{\partial}_{z_1} \bar{\partial}_{\bar{z}_2} Q_k(z_1, z_2) .
$$

and that the number variance for zeros has the asymptotics,

$$
Var(\mathcal{N}_U) = k^{1/2} \left[\nu_1 Vol_1(\partial U) + O(k^{-\frac{1}{2} + \varepsilon}) \right],
$$

where $\nu_1 = \frac{\zeta(3/2)}{8\pi^{3/2}}$. Thus, the same formula is valid in the limit $t \to \infty$ of heat kernel random metrics. Similar formulae for variances of X_{U} and X_{U} heat kernel random metrics. Similar formulae for variances of X_U and X_{ϕ} for random metrics can be derived from the explicit formula (9) for $I_{2,k}(t, \rho)$ for any t in the heat kernel measure ensembles. The details are lengthy and will be presented elsewhere.

1.3. Asymptotic central limit theorem

It is shown in [18] that the fluctuations of the smooth linear statistics X_f in the case of random zeros tend to a Gaussian field with variance $\mathcal{N}(0, \sqrt{\kappa_1} || \partial \bar{\partial} f ||_2)$, see §5.2 for more details. This result holds when $t = \infty$ for heat kemple proders metrics. The englocaus results for (2) do not seem for heat kernel random metrics. The analogous results for (3) do not seem to be known at present.

It is very plausible that for general times t_k the smooth linear statistics X_f with respect to the heat kernel measure $d\mu_k^{t_k}$ are also asymptotically normal, with a related variance. Such an asymptotic central limit theorem would be a concrete measure of how closely heat kernel random metrics compare to random singular metrics defined by point processes of random zero sets, or to those studied in $[1, 5, 14, 20]$, where the fluctuations of linear statistics of eigenvalues tend to a Gaussian free field. We plan to investigate

²In the notations of [17], $I_{2,k}$ corresponds to $4\pi^2 Q_k$.

the asymptotic normality of fluctuations of random metrics in future work. These asymptotic normality results would also give a comparison of heat kernel random metrics to Liouville random metric. The fluctuations in the latter case are of the type of Gaussian multiplicative chaos.

1.4. Discussion

Heat kernel random metrics are the metrics obtained by starting at the background metric ω_{ϕ_I} and following a Brownian motion on \mathcal{P}_{N_k} for time t.

The formula (9) reflects the geometry of Brownian motion of the nonpositively curved symmetric space \mathcal{P}_{N_k} , which is very different from that of Euclidean space, see §3 for background. First, due to non-isotropic nature of the Haar measure, the heat measure is concentrated along the $SU(N_k)$ -orbit of a distinguished element δ_{N_k} , the half-sum of the positive roots. Second, in the radial direction the heat kernel measure concentrates in a kind of annulus of radius t around the $SU(N_k)$ -orbit of δ_{N_k} , see §3.2. This is why, as $t \to \infty$, the heat kernel measure becomes supported on the ideal boundary $\partial_{\infty}P_{N_k}$. Its $SU(N_k)$ invariance implies that the boundary measure is the same as the measure on zero sets of holomorphic sections used in [17].

As mentioned above, 'heat kernel' random metrics are not like the random metrics of Liouville quantum gravity. On a very heuristic level, one may understand the difference by thinking of $SU(N_k)$ -invariance as a discretization of invariance under the group $SDiff_{\omega_0}(M)$ of symplectic diffeomorphisms of the background symplectic form (M, ω_0) . This is far from the invariance group of Liouville theory. It is plausible that the only $SDiff_{\omega_0}(M)$ invariant probability measure on $\mathcal{K}_{[\omega_0]}$ is δ_{ω_0} .

This concentration of measure phenomenon, that heat kernel random metrics concentrate on the background, is the same phenomenon that occurs for random zeros in [17, 18] and for the quantum Hall point process in [14]. Random zero sets of N zeros or random point configurations of N electrons also concentrate at the background metric as $N \to \infty$. In these cases it is customary to dilate the space to obtain configurations of constant density. In the case of random metrics, if one dilates small balls of radius $\frac{1}{\sqrt{k}}$ around a point z_0 by the factor \sqrt{k} , then the random metrics become metrics on \mathbb{C}
and the 2 point completion function of the limit measure acquired the perm and the 2-point correlation function of the limit measure acquires the new term $\delta(z-w)$.

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2. Bergman metrics

We briefly review the properties of Bergman metrics, referring to [8, 16] for further background. As above, (M, ω, J) can be any compact Kähler manifold with integral Kähler form. The Riemannian metric is $g(X, Y) =$ $\omega(JX, Y)$. The simplest case is that of a Riemann surface, where a Kähler class is the same as a conformal class of metrics with fixed area. Instead of parametrizing metrics $g = e^u g_0$ by the Liouville field u we parametrize them by the Kähler potential ϕ , i.e. $\omega_{\phi} = \omega_0 + i\partial\bar{\partial}\phi$, relative to the reference Kähler form ω_0 .

Bergman metrics of degree k are special Kähler metrics induced by holomorphic embeddings

$$
\iota_s(z) = [s_1, \ldots, s_{N_k}] : M \to \mathbb{CP}^{N_k - 1}
$$

of M into complex projective space. Here, $\{s_i\}$ is a basis of the space $H^0(M,L^k)$ of holomorphic sections of powers L^k of an ample line bundle $L \to M$ with first Chern class $c_1(L)=[\omega_0]$. Also $N_k = \dim H^0(M,L^k)$. Given a reference basis $\{s_i\}$ one obtains all others by applying an element $A \in GL(N_k, \mathbb{C})$ to it $s^A{}_j = \sum A_{jl} s_l$ and induces the embedding

$$
\iota_{s_A}:M\to\mathbb{CP}^{N_k-1},\ \ \iota_{s_A}=A\circ\iota_s.
$$

The associated Bergman metric is then,

(13)
$$
\iota_{s_A}^* \omega_{FS} = \frac{1}{k} i \partial \bar{\partial} \log \sum_{j=1}^{N_k} |s^A{}_j(z)|^2.
$$

Since $U(N_k)$ is the isometry group of ω_{FS} , the space of metrics is the quotient symmetric space $\mathcal{P}_{N_k} = GL(N_k,\mathbb{C})/U(N_k)$. With no loss of generality one may restrict to $SL(N_k, \mathbb{C})$ and obtain the quotient $SL(N_k, \mathbb{C})/SU(N_k)$.

We choose a basis of sections $\{s_i(z)\} = \{s_1(z), ..., s_{N_k}(z)\}\$ of $H^0(M, L^k)$ which is orthonormal with respect to the reference (background) metric h_0^k on L^k and the corresponding Kähler metric $\omega_0 = -\frac{1}{k} i \partial \bar{\partial} \log h_0^k$ on M

(14)
$$
\frac{1}{V} \int_M \bar{s}_i(z) s_j(z) h_0^k \frac{\omega_0^n}{n!} = \delta_{ij},
$$

where $n = \dim M$. The Bergman kernel of the background metric is the kernel of the orthogonal projection onto $H^0(M,L^k)$ with respect to the inner product above, and is given by

(15)
$$
B_k(z_1, z_2) = \sum_{j=1}^{N_k} s_j(z_1) \bar{s}_j(z_2)
$$

Given a positive Hermitian matrix $P = P_{ij}$ the associated Bergman metric is,

(16)
$$
\omega_{a\bar{b}}(z) = \frac{1}{k} \partial_a \bar{\partial}_{\bar{b}} \log \bar{s}_i(z) P_{ij} s_j(z).
$$

In terms of $A \in GL(N_k, \mathbb{C})$ above, $P = A^{\dagger} A$. We introduce the Bergman potential as follows

(17)
$$
\phi_P = \frac{1}{k} \log \bar{s}_i(z) P_{ij} s_j(z) = \frac{1}{k} \log |\langle e^{\Lambda} U s(z), U s(z) \rangle|^2.
$$

A key property is that $\mathcal{K}_{[\omega_0]} = \bigcup_k \mathcal{B}_k$, i.e. the full space of metrics in the kishler close is the closure of the set of Bergman metrics. Hence a fixed Kähler class is the closure of the set of Bergman metrics. Hence $\mathcal{K}_{[\omega_0]}$ is well approximated by \mathcal{B}_k for large k, and there are now many results showing that it is well approximated in much stronger geometric ways. This approximation problem was raised by S. T. Yau in [21], see [16] for background.

2.1. Berezin kernel

The key invariant is the Berezin kernel (5), given in the above notation by

(18)
$$
\rho = \frac{|\langle s(z_1), s(z_2) \rangle|^2}{|s(z_1)|^2 |s(z_2)|^2},
$$

or in terms of the Bergman kernel

(19)
$$
\rho = P_k^2(z_1, z_2) := \frac{|B_k(z_1, z_2)|^2}{B_k(z_1, z_1)B_k(z_2, z_2)}.
$$

2.2. Matrix-metric correspondence

The matrix-metric correspondence Eq. (16) uses a choice of basis $\{s_i\}$ of $H^0(M, L^k)$. Any natural measure on \mathcal{B}_k must be independent of the choice of this basis. We pause to describe such natural measures.

Any Kähler metric $\omega = \omega_0 + i\partial\bar{\partial}\phi$ in $\mathcal{K}_{[\omega_0]}$ induces an inner product $Hilb_k(\phi)$ on $H^0(M,L^k)$ by the rule

$$
\langle s_1, s_2 \rangle_{\mathrm{Hilb}_k(\phi)} = \int_M \bar{s}_1(\bar{z}) s_2(z) h^k \frac{\omega^n}{n!}.
$$

Given a background inner product $G_0 = Hilb_k(\phi_0)$, any other inner product has the form $\langle s_1, s_2 \rangle_G = \langle P_G s_1, s_2 \rangle_{G_0}$ where P_G is a positive Hermitian operator on $H^0(M,L^k)$ with respect to G_0 . It has a well-defined polar decomposition $e^{\Lambda}U$ where $U \in U(G_0)$ is unitary with respect to G_0 . Its eigenvalues are encoded by the diagonal matrix Λ_G and its eigenvectors are encoded by U .

In making calculations, we need to parametrize such positive Hermitian operators by positive Hermitian matrices, which requires a choice of a G_0 orthonormal basis of $H^0(M,L^k)$. Any measure intrinsically defined on the space of positive Hermitian operators will be independent of the choice of basis. Haar measure and the heat kernel are examples of such measures.

3. Heat kernel

In this section we review the heat kernel on \mathcal{P}_N^3 . Bergman metrics are unchanged if the positive Hermitian matrix P is multiplied by a scalar, so we may normalize P so that det $P = 1$. Then $\mathcal{P}_N = G/K$ where $G = SL(N, \mathbb{C})$ and $K = SU(N)$. We denote by ℓ the Lie algebra of the maximal compact subgroup $K \subset G$ and let $\mathfrak{g} = \mathfrak{k} \oplus \mathfrak{p}$. Let \mathfrak{a} be a maximal abelian subspace of **p** and let $\ell = \dim \mathfrak{a}$. The set of positive roots is denoted by R_+ . The roots are $e_i - e_j$, and the positive roots satisfy $i < j$ and have multiplicity one. For $SL(N,\mathbb{C})/SU(N)$ the half sum of the positive roots is the element $\delta_N = \left(-\frac{N-1}{2}, -\frac{N-3}{2}, \ldots, \frac{N-1}{2}\right)$. For background, see [12, 13].
We refer to the matrix decomposition $P = U^{\dagger}e^{\Delta}U$ for Λ .

We refer to the matrix decomposition $P = U^{\dagger}e^{\Lambda}U$ for $\Lambda = \text{diag}(\lambda_1, \ldots,$ λ_N , and $U \in U(N)$ as 'polar coordinates' on \mathcal{P}_N , where real numbers $\lambda_j \in$ $(-\infty, +\infty)$ correspond to the Cartan elements of $SL(N, \mathbb{C})$.

³In this section we adopt shorthand notation $N = N_k$.

The CK (Cartan-Killing) metric is given by

$$
(20) \qquad \qquad ds^2 = \text{Tr}(P^{-1}dP)^2
$$

for $P \in GL(N, \mathbb{C})/U(N)$. This metric is bi-invariant under the action of $GL(N,\mathbb{C}).$

The associated volume form dV on the symmetric space $SL(N,\mathbb{C})/$ $SU(N)$ of positive Hermitian matrices with det $P=1$ is the bi-invariant Haar measure,

(21)
$$
dV = \delta \left(\sum_{j=1}^{N} \lambda_j \right) \Delta^2(e^{\lambda}) \prod_{j=1}^{N} d\lambda_j \cdot \frac{[dU]}{[dU_{U(1)^N}]},
$$

where $\left|\frac{dU}{dt}\right|$ is the standard Haar measure on unitary group.

3.1. Heat kernel measure on $SL(N, \mathbb{C})/SU(N)$

Following Gangolli [10] (Proposition 3.2; see also [3], section 2), the heat kernel on $SL(N,\mathbb{C})/SU(N)$ with respect to the standard CK (Cartan-Killing) metric is given in 'polar coordinates' (λ, U) on \mathcal{P}_N by

(22)
$$
d\mu_t = g_t(\lambda)dV = C(t, N)\frac{\Delta(\lambda)}{\Delta(e^{\lambda})}e^{-\frac{1}{4t}\sum_{j=1}^N\lambda_j^2}dV.
$$

Here, $\Delta(\lambda) = \prod_{i < j} (\lambda_j - \lambda_i)$ is the standard Vandermonde determinant.

The normalization constant $C(t, N)$ in (22) is fixed by the condition that μ_t is the probability measure $\int d\mu_t = 1$,

(23)
$$
C(t,N) = \frac{\sqrt{N}}{2\pi(\sqrt{4\pi t})^{N^2-1}}e^{-\frac{t}{12}N(N^2-1)}.
$$

In deriving this we use the volume of the unitary group

$$
Vol U(N) = (2\pi)^{N(N+1)/2} / \prod_{j=1}^{N} j!,
$$

see e.g. [15]. The factor $e^{-\frac{t}{12}N(N^2-1)}$ is $e^{-t||\delta_N||^2}$ and arises because $||\delta_N||^2$ is the bottom of the spectrum of the Laplacian. Putting (21) and (23) together, we get the following expression

(24)
$$
d\mu_t = \frac{\sqrt{N}}{2\pi(\sqrt{4\pi t})^{N^2 - 1}} e^{-\frac{t}{12}N(N^2 - 1)} \delta\left(\sum_{j=1}^N \lambda_j\right) \Delta(\lambda)
$$

$$
\Delta(e^{\lambda}) e^{-\frac{1}{4t} \sum_{j=1}^N \lambda_j^2} \prod_{j=1}^N d\lambda_j \cdot \frac{[dU]}{[dU_{U(1)^N}]}
$$

for the heat kernel measure on $SL(N,\mathbb{C})/SU(N)$ with respect to the CK metric.

3.2. Geometry of the heat kernel and Brownian motion

In [2] it is proved that the mass of the heat kernel concentrates along the exponential image of the $U(N)$ -orbit of the δ_N -axis in a small annulus centered at $2|\delta_N|t$.

If we write $H = \text{diag}(\lambda)$, then the Gaussian factor $t^{-(N^2-1)/2}e^{-\frac{\|H\|^2}{4t}}$ is similar to the heat kernel of Euclidean space. But this Gaussian factor must compete with the exponential volume growth factor $\Delta(e^{\lambda})$ and the factor $e^{-\frac{t}{12}N(N^2-1)}$ due to the existence of a spectral gap for Δ . The wellknown factor $\Delta(\lambda)$ pushes the eigenvalues of $\log P$ apart. The factor $J(H)$ is bounded by $e^{2\langle \delta_N, \vec{\lambda} \rangle}$ and a simplified expression for the heat kernel is $e^{-t|\delta_N|^2 + \langle \lambda, \rho_N \rangle - \frac{|\lambda|^2}{2t}}$. The maximum of the exponent occurs when $\vec{\lambda} = 2t\delta_N$.

Following [2], let $\gamma(t)$ be a positive function with $\sqrt{t}\gamma(t) \to \infty$ as $t \to \infty$, Following [2], let $\gamma(t)$ be a positive function with $\sqrt{t} \gamma(t) \to \infty$ as $t \to \infty$,
and let $R(t)$ be a positive function such that $R(t)/\sqrt{t} \to \infty$. Consider the annulus

$$
A(2|\delta_N|t - R(t), 2|\delta_N|t + R(t))
$$

:= $\{H : 2|\delta_N|t - R(t) \le |H| \le 2|\delta_N|t + R(t)\} \subset \mathfrak{a}$

and consider the solid cone

$$
\Gamma(t)
$$
 = solid cone around the δ_N axis of angle $\gamma(t)$,

and let

$$
\Omega(t) = A(2|\delta_N|t - R(t), 2|\rho|t + R(t)) \cap \Gamma(t).
$$

Then, according to Theorem 1 of [2],

(25)
$$
\int_{U(N) \exp \Omega(t)U(N)} d\mu_t \to 1, \ \ t \to \infty.
$$

A Brownian motion proof of this result is given in [4]. It shows that as $t \to \infty$ the mass of μ_t moves off to a component of the ideal boundary (at infinity) of \mathcal{P}_N . In §5 we discuss this boundary.

3.3. Scaling and dilation

In the large k limit, the symmetric space metric on \mathcal{B}_k , when properly scaled, tends to the Mabuchi metric g_M on $\mathcal{K}_{[\omega_0]}$ (see [6]). The Mabuchi distance function is induced by the Riemannian metric on $\mathcal{K}_{[\omega_0]}$ defined by $||\delta \phi||_{\phi_0}^2 = \int_M (\delta \phi)^2 \omega_{\phi}^n/n!$ where $\omega_{\phi} = \omega_0 + i \partial \bar{\partial} \phi$. We refer to [16] for background. For $\mathcal{L}_{M}(\delta\phi)^{2}\omega_{\phi}^{n}/n!$ where $\omega_{\phi} = \omega_{0} + i\partial\bar{\partial}\phi$. We refer to [16] for background. For all k, $\mathcal{B}_k \subset \mathcal{K}_{[\omega_0]}$. If we rescale the CK metric $g_{CK,k}$ (20) as $g_k = \varepsilon_k^2 g_{CK,k}$, with $\varepsilon_k = k^{-1} N_k^{-1/2}$, then $g_k \to g_M$ on $T \mathcal{B}_k$. Thus, a ball of radius one with respect to the usual CK metric $g_{CK,k}$ has radius approximately ε_k with respect to the Mabuchi distance. It is obviously desirable to consider the heat kernel measures for this rescaled sequence of metrics.

If we rescale the CK metric to $g_k = \varepsilon_k^2 g_{CK,k}$ the corresponding Laplacian scales as $\Delta_{g_k} \to \varepsilon_k^{-2} \Delta_{g_{CK,k}}$. It follows that the heat operator scales as

$$
\exp t\Delta_{g_k} = \exp t \varepsilon_k^{-2} \Delta_{g_{CK,k}}.
$$

In effect, it is only the time that is rescaled and the rescaled heat kernel is $p_k(\varepsilon_k^{-2}t, I, P).$

4. One and two point correlation functions of random metrics

In this section, we calculate the one and two point functions of the random Kähler potential; in the introduction, the latter was stated to be (9) . We use the notation $\mathbf{E} = \mathbf{E}_k$ for the expectation, which at the beginning could be with respect to any $U(N_k)$ -invariant measure and then specializes to the heat kernel measures. For simplicity of notation we often abbreviate N_k by N and drop the explicit k-dependence.

As mentioned in the introduction, the one and two-point functions are the data required to study the mean and variance of the area random variables X_U . Evidently,

$$
\mathbf{E}_k X_U = \int_U \mathbf{E}_k \,\omega,
$$

\n
$$
\mathbf{Var}(X_U) = \int_{U \times U} \mathbf{E}_k [\omega(z_1) \omega(z_2)] - \int_{U \times U} \mathbf{E}_k [\omega(z_1)] \mathbf{E}_k [\omega(z_2)].
$$

The integrands are the one- and two-point correlation functions.

4.1. One point function of the Kähler potential

The argument here follows [9]. Using the integral representation of the logarithm

(26)
$$
\log \alpha = \frac{1}{\tau} + \gamma - \int_0^\infty x^{\tau - 1} e^{-\alpha x} dx + \mathcal{O}(\tau),
$$

where γ is the Euler constant, we can rewrite the expectation value of the Kähler potential ϕ_P (17) relative to the Kähler potential ϕ_I of the background Bergman metric as

(27)
$$
\mathbf{E}_{k}[\phi_{P}(z) - \phi_{I}(z)]
$$

\n
$$
= \mathbf{E}_{k} \left[\frac{1}{k} \log \frac{\bar{s}(z)U^{\dagger}e^{\Lambda}Us(z)}{|s(z)|^{2}} \right]
$$

\n
$$
= \frac{1}{k} \lim_{\tau \to 0} \frac{1}{\tau} + \gamma - \int_{0}^{\infty} x^{\tau-1} dx \int_{\mathcal{P}_{N}} e^{-\text{Tr} e^{\Lambda}U\Psi U^{\dagger}} d\mu_{t},
$$

where introduced the matrix $\Psi_{jl} = x s_j(z) \bar{s}_l(z) / |s(z)|^2$. The integration over the unitary group can be carried out using Harish-Chandra-Itzykson-Zuber formula, see e. g. [15, 22] for background. Namely, for any two Hermitian matrices A and B with eigenvalues a_j and b_j

(28)
$$
\int_{U(N)} \frac{[dU]}{\text{Vol } U(N)} \exp \left(\mu \text{Tr} A U B U^{\dagger}\right)
$$

$$
= \left(\prod_{p=1}^{N-1} p! \right) \mu^{-N(N-1)/2} \frac{\det \left(e^{\mu a_j b_l}\right)_{1 \le j,l \le N}}{\Delta(a) \Delta(b)}.
$$

It is not hard to check that this expression is well defined even if some of the eigenvalues coincide. This is the case for the matrix Ψ which has $N-1$ zero eigenvalues and one non-zero eigenvalue equal to x . Hence the integral on the right hand side of (27) is z-independent and we immediately conclude that the expectation value of the Bergman metric is equal to the background Bergman metric

$$
\mathbf{E}_k[\omega_{a\bar{b}}] = \omega_{\phi_I, a\bar{b}}
$$

In fact, this is true for any eigenvalue-type measure [9], since the HCIZ integral depends on eigenvalues only. Note, that so far no assumptions on k have been made in this calculation. Considering now the limit of k large, we invoke the Bergman kernel expansion [9] to show that the background Bergman metric ω_{ϕ_I} tends to the reference Kähler metric,

$$
\omega_{\phi_I, a\bar{b}} = \omega_{0\, a\bar{b}} + \mathcal{O}(1/k).
$$

Now we would like to consider the variance of X_U from its mean $\omega_0(U)$.

4.2. The two point function

In this section we prove the formula (9) for the two point function, and then discuss its asymptotics in various regimes. In the terminology of [17] we are finding a 'bi-potential' for the variance. Although the calculations of this bi-potential are completely different from the case of random holomorphic sections in [17], the final formulae are somewhat similar and when $t \to \infty$ they are identical.

Recall that the two-point function (4) of the Kähler potential is the sum of the background term $\phi_I(z)\phi_I(w)$ plus the variance term $I_{2,k}(\rho)$. Instead of using the formula (8) for this term, we take the approach of writing

(29)
$$
K_{2,k}(z, w) - \phi_I(z)\phi_I(w)
$$

= $\frac{1}{k^2} \mathbf{E}_k \left[\log \frac{\bar{s}(z_1)U^{\dagger}e^{\Lambda}Us(z_1)}{|s(z_1)|^2} \log \frac{\bar{s}(z_2)U^{\dagger}e^{\Lambda}Us(z_2)}{|s(z_2)|^2} \right]$
= $\frac{1}{k^2} \lim_{\tau_1, \tau_2 \to 0} (I_{2,k}(t, \rho, \tau_1, \tau_2) + \rho$ -independent terms)

where we do not write down the ρ -independent terms, since ultimately we are interested in the dependence of the correlation function on coordinates, which enter only through $\rho(z_1, z_2)$. Here

(30)
$$
I_{2,k}(t,\rho,\tau_1,\tau_2) = \iint_0^\infty x_1^{\tau_1-1} x_2^{\tau_2-1} dx_1 dx_2 \int_{\mathcal{P}_N} e^{-\text{Tr}\,e^{\Lambda}U\Phi U^\dagger} d\mu_t.
$$

and we introduced the matrix $\Phi_{jl} = x_1 \frac{s_j(z_1)\bar{s}_l(z_1)}{|s(z_1)|^2} + x_2 \frac{s_j(z_2)\bar{s}_l(z_2)}{|s(z_2)|^2}$. It has rank 2 with two non-zero eigenvalues given by

$$
\phi_{1,2} = \frac{1}{2} \left(x_1 + x_2 \pm \sqrt{(1 - \rho)(x_1 - x_2)^2 + \rho(x_1 + x_2)^2} \right).
$$

Applying the HCIZ formula (28) to the unitary integration in Eq. (30) we obtain

$$
(-1)^{N(N-1)/2} \frac{N!(N-1)!}{(\phi_1 \phi_2)^{N-2}(\phi_1 - \phi_2)} \frac{e^{-\phi_1 e^{\lambda_1} - \phi_2 e^{\lambda_2}}}{\prod_{j=2}^N (e^{\lambda_1} - e^{\lambda_j}) \prod_{l=3}^N (e^{\lambda_2} - e^{\lambda_l})}
$$

where we used the fact that the integration measure is symmetric in eigenvalues λ 's.

Now we use the explicit form of the heat kernel measures μ_t (24). The integral $I_{2,k}(t, \rho, \tau_1, \tau_2)$ in Eq.(29) with this eigenvalue measure can be written as

$$
I_{2,k}(t, \rho, \tau_1, \tau_2) = (-1)^{N(N-1)/2} C(t, N) \frac{\text{Vol } U(N)}{(2\pi)^N} N!(N-1)!
$$

$$
\iint_0^\infty \frac{x_1^{\tau_1 - 1} x_2^{\tau_2 - 1} dx_1 dx_2}{(\phi_1 \phi_2)^{N-2} (\phi_1 - \phi_2)}
$$

$$
\int_{-\infty}^\infty dy \int \prod_{j=1}^N d\lambda_j \ \Delta(\lambda) \Delta_{12}(e^{\lambda})
$$

$$
\exp\left(-\frac{1}{4t} \sum_{j=1}^N \lambda_j^2 + iy \sum_{j=1}^N \lambda_j - \phi_1 e^{\lambda_1} - \phi_2 e^{\lambda_2}\right)
$$

where we defined the partial Vandermonde determinant $\Delta_{12}(e^{\lambda}) = \prod_{3 \leq j < l \leq N} (e^{\lambda_j} - e^{\lambda_l})$, which excludes the first two eigenvalues e^{λ_1} and e^{λ_2} .
The *u*-integration enforces the delta-function constra The y-integration enforces the delta-function constraint in the measure (24) .

 $\,$

Using antisymmetry of $\Delta(\lambda)$ under exchange of two eigenvalues, $\Delta_{12}(e^{\lambda})$ can be replaced by $(-1)^{1+N(N-1)/2}(N-2)!e^{\sum_{l=3}^{N}(l-3)\lambda_l}$ inside the integral, which leads to the further simplification

(31)
$$
I_{2,k}(t, \rho, \tau_1, \tau_2)
$$

= $-\frac{\text{Vol }U(N)}{(2\pi)^N}C(t, N)N!(N-1)!(N-2)!$

$$
\iint_0^\infty \frac{x_1^{\tau_1-1}x_2^{\tau_2-1}dx_1dx_2}{(\phi_1\phi_2)^{N-2}(\phi_1-\phi_2)}.
$$

$$
\int_{-\infty}^\infty dy \int \prod_{j=1}^N d\lambda_j \Delta(\lambda) e^{\sum_{j=1}^N \left(-\frac{1}{4t}\lambda_j^2 + iy\lambda_j\right) + \sum_{l=3}^N (l-3)\lambda_l - \phi_1 e^{\lambda_1} - \phi_2 e^{\lambda_2}},
$$

Thus after the HCIZ integration we got rid of most difficult factor $\Delta(e^{\lambda})$ and left with a Gaussian integral with a polynomial amplitude, except for the terms $\phi_1 e^{\lambda_1} + \phi_2 e^{\lambda_2}$ in the exponent. Note that the k-dependence is entirely in the variable ρ inside ϕ_1, ϕ_2 .

The next step is to calculate the integrals in the last line of (31),

(32)
$$
\mathcal{I}_{2,k}(t, \phi_1, \phi_2) := \int_{-\infty}^{\infty} dy \int \prod_{j=1}^{N} d\lambda_j \, \Delta(\lambda)
$$

$$
e^{\sum_{j=1}^{N} \left(-\frac{1}{4t} \lambda_j^2 + iy\lambda_j\right) + \sum_{l=3}^{N} (l-3)\lambda_l - \phi_1 e^{\lambda_1} - \phi_2 e^{\lambda_2}}.
$$

Our strategy is to Taylor-expand the exponent of (32) in powers of e^{λ} ,

$$
e^{-\phi_1 e^{\lambda_1} - \phi_2 e^{\lambda_2}} = \sum_{m,n=0}^{\infty} \frac{(-\phi_1)^n (-\phi_2)^m}{n! m!} e^{n\lambda_1 + m\lambda_2},
$$

then perform gaussian integration in λ term-by-term, and finally re-sum the resulting series, i. e. un-do the Taylor expansion. We use the identity

(33)
\n
$$
\prod_{j=1}^{N} \left(\int_{-\infty}^{\infty} d\lambda_j \right) \Delta(\lambda) e^{\sum_{j=1}^{N} \left(-\frac{1}{4t} \lambda_j^2 + \mu_j \lambda_j \right)}
$$
\n
$$
= (2\pi)^{N/2} (2t)^{N^2/2} \Delta(\mu) e^{t \sum_{j=1}^{N} \mu_j^2}
$$

to compute the eigenvalue integral and get

(34)
$$
\mathcal{I}_{2,k}(t,\phi_1,\phi_2) = \sum_{n,m=0}^{\infty} \frac{(n-m)}{n!m!}(-\phi_1)^n(-\phi_2)^m \prod_{l=0}^{N-3} (n-l)(m-l)
$$

$$
\int_{-\infty}^{\infty} dy \, e^{t(n+iy)^2 + t(m+iy)^2 + t \sum_{l=0}^{N-3} (l+iy)^2}.
$$

Due to the factor $\prod_{l=0}^{N-3} (n-l)(m-l)$, all terms with $m, n < N-2$ have
coefficient zero, so we can shift summation indices $n \to n - (N-2)$, $m \to$ coefficient zero, so we can shift summation indices $n \to n - (N - 2)$, $m \to$ $m - (N - 2)$. Integrating over y in (34) and plugging the result back to (31) we obtain

(35)
$$
I_{2,k}(t, \rho, \tau_1, \tau_2)
$$

= $-e^{-t(N-1)^2/N} \int \int_0^\infty \frac{x_1^{\tau_1-1} x_2^{\tau_2-1} dx_1 dx_2}{\phi_1 - \phi_2} (\phi_1 \partial_{\phi_1} - \phi_2 \partial_{\phi_2})$

$$
\sum_{n,m=0}^\infty \frac{(-\phi_1)^n (-\phi_2)^m}{n!m!} e^{\frac{t}{2} \frac{N-2}{N} (n+m)^2 + \frac{t}{2} (n-m)^2 + t \frac{(N-1)(N-2)}{N} (n+m)}
$$

.

Now we can re-sum the series using the identity

$$
\frac{1}{2\pi t} \int_{-\infty}^{\infty} d\lambda_1 d\lambda_2 e^{-\frac{1}{2t}(\lambda_1^2 + \lambda_2^2) - \phi_1 e^{a\lambda_1 + \lambda_2} - \phi_2 e^{a\lambda_1 - \lambda_2}}
$$
\n
$$
= \sum_{n,m=0}^{\infty} \frac{(-\phi_1)^n (-\phi_2)^m}{n!m!} e^{\frac{t}{2}a^2(n+m)^2 + \frac{t}{2}(n-m)^2}
$$

Replacing the series in (35) by the integral, and changing variables $x_1 \rightarrow e^{-a\lambda_1 - t\frac{(N-1)(N-2)}{N}} x_1, x_2 \rightarrow e^{-a\lambda_1 - t\frac{(N-1)(N-2)}{N}} x_2$ makes it possible to carry out the gauge integration in λ giving the gaussian integration in λ_1 , giving

$$
I_{2,k}(t, \rho, \tau_1, \tau_2) = -\frac{e^{-t/2}}{\sqrt{2\pi t}} \iint_0^\infty \frac{x_1^{\tau_1 - 1} x_2^{\tau_2 - 1} dx_1 dx_2}{\phi_1 - \phi_2} \left(\phi_1 \partial_{\phi_1} - \phi_2 \partial_{\phi_2}\right) \int_{-\infty}^\infty d\lambda \, e^{-\frac{1}{2t} \lambda^2 - \phi_1 e^{\lambda} - \phi_2 e^{-\lambda}}
$$

Interestingly, the factor $e^{-t|\delta_N|^2}$ coming from the spectral gap has now disappeared from the formula. It follows from the integral representation Eq. (26), that the singular in $\tau_{1,2}$ terms in (29) are ρ -independent. Therefore after taking the derivative of $I_{2,k}(t, \rho, \tau_1, \tau_2)$ with respect to ρ , we can set $\tau_1 = \tau_2 = 0$. Since the measures μ_t depend on t we henceforth denote the corresponding term in the two point function by $I_{2,k}(t, \rho)$. Using

$$
\partial_{\rho}\phi_{1,2} = \pm \frac{x_1 x_2}{\phi_1 - \phi_2}, \quad \partial_{\rho}\frac{1}{\phi_1 - \phi_2} = -\frac{2x_1 x_2}{(\phi_1 - \phi_2)^3}
$$

we get

$$
\partial_{\rho}I_{2,k}(t,\rho) = \frac{2e^{-t/2}}{\sqrt{2\pi t}} \iint_0^{\infty} dx_1 dx_2 \int_{-\infty}^{\infty} d\lambda \, e^{-\frac{1}{2t}\lambda^2 - \phi_1 e^{\lambda} - \phi_2 e^{-\lambda}}
$$

$$
\frac{\sinh \lambda}{(\phi_1 - \phi_2)^3} (\phi_1 + \phi_2 + (\phi_1 - \phi_2)(\phi_1 e^{\lambda} - \phi_2 e^{-\lambda}))
$$

Now we introduce new coordinates (r, θ) as

$$
r \cos \theta = \sqrt{\rho} (x_1 + x_2), \quad r \sin \theta = \sqrt{1 - \rho} (x_2 - x_1),
$$

with the range

$$
r \in [0, \infty)
$$
, $\theta \in [-\alpha, +\alpha]$ where $\cos \alpha = \sqrt{\rho}$.

In terms of r and θ the integrals can be written as

$$
\partial_{\rho}I_{2,k}(t,\rho) = -\frac{e^{-t/2}}{\sqrt{2\pi t}} \frac{1}{\sqrt{\rho(1-\rho)}}
$$

$$
\int_{0}^{\infty} r dr \int_{-\alpha}^{\alpha} d\theta \int_{-\infty}^{\infty} d\lambda \, e^{-\frac{1}{2t}\lambda^{2}-r\left(\frac{\cos\theta}{\sqrt{\rho}}\cosh\lambda+\sinh\lambda\right)}
$$

$$
\left[\frac{\cos\theta}{r^{2}\sqrt{\rho}}+\frac{1}{2r}\left(\frac{\cos\theta+\sqrt{\rho}}{\sqrt{\rho}}e^{\lambda}-\frac{\cos\theta-\sqrt{\rho}}{\sqrt{\rho}}e^{-\lambda}\right)\right] \sinh\lambda.
$$

Introducing new variable $x\sqrt{\rho} = \cos\theta$ and rearranging exponents, we get

$$
\partial_{\rho}I_{2,k}(t,\rho) = -\frac{e^{-t/2}}{\sqrt{2\pi t}} \frac{1}{\sqrt{1-\rho}} \int_0^{\infty} r dr \int_1^{1/\sqrt{\rho}} \frac{dx}{\sqrt{1-\rho x^2}} \int_{-\infty}^{\infty} d\lambda \, e^{-\frac{1}{2t}\lambda^2 + \lambda}
$$

$$
e^{-rx \cosh\lambda} \left[\frac{x}{r^2} \left(e^{-r \sinh\lambda} - e^{r \sinh\lambda} \right) + \frac{\sinh\lambda}{r} \left((x+1)e^{-r \sinh\lambda} + (x-1)e^{r \sinh\lambda} \right) \right].
$$

Integrating over r we obtain

$$
\partial_{\rho}I_{2,k}(t,\rho) = -\frac{e^{-t/2}}{\sqrt{2\pi t}} \frac{1}{\sqrt{1-\rho}} \int_{1}^{1/\sqrt{\rho}} \frac{dx}{\sqrt{1-\rho x^{2}}} \int_{-\infty}^{\infty} d\lambda \, e^{-\frac{1}{2t}\lambda^{2}+\lambda}
$$

$$
\left[x \log \left(\frac{x \cosh \lambda - \sinh \lambda}{x \cosh \lambda + \sinh \lambda}\right) + 2 \sinh \lambda \frac{x^{2} \cosh \lambda - \sinh \lambda}{x^{2} \cosh^{2} \lambda - \sinh^{2} \lambda}\right].
$$

After integrating the log term by parts, we can perform the x integration

(36)
$$
\partial_{\rho}I_{2,k}(t,\rho) = \frac{2t}{\rho} - \frac{2e^{-t/2}}{\sqrt{2\pi t}} \frac{\sqrt{1-\rho}}{\rho} \int_{-\infty}^{\infty} d\lambda
$$

$$
\int_{1}^{1/\sqrt{\rho}} \frac{dx}{\sqrt{1-\rho x^{2}}} \cdot \frac{e^{-\frac{1}{2t}\lambda^{2}}\cosh\lambda\sinh^{2}\lambda}{x^{2}\cosh^{2}\lambda - \sinh^{2}\lambda}
$$

$$
= \frac{2t}{\rho} - \frac{e^{-t/2}}{\sqrt{2\pi t}} \frac{\sqrt{1-\rho}}{\rho} \int_{-\infty}^{\infty} d\lambda \, e^{-\frac{1}{2t}\lambda^{2}} \frac{\cosh\lambda}{\sqrt{\coth^{2}\lambda-\rho}}
$$

$$
\log \frac{\sqrt{\coth^{2}\lambda-\rho}+\sqrt{1-\rho}}{\sqrt{\coth^{2}\lambda-\rho}-\sqrt{1-\rho}}.
$$

This completes the calculation of (9).

An important application of these asymptotics is to calculate the variances of the linear statistics X_U (3) and its smooth analogue X_f . This can be done precisely as in Section 4 of [17] and Section 3 of [18], but substituting the formula for $I_{2,k}(t,\rho)$ for Q_k . The details are rather lengthy and will be presented elsewhere.

The remainder of the article is devoted to the asymptotics of $I_{2,k}(t, \rho)$ as $k \to \infty$ with various relations between t and k. To understand the various regimes, it should be recalled that the metric rescaling in §3.3 is necessary to ensure that balls in the Cartan-Killing metric on \mathcal{B}_k maintain their size with respect to the limiting Mabuchi metric on $\mathcal{K}_{[\omega_0]}$. Without rescaling, the Brownian motion relative to the CK metric is probing metrics only at distance $d_k = \varepsilon_k^2 t$ from the initial background metric with respect to the limiting metric.

In all of the regimes, the key to finding the scaling asymptotics is to work out the behavior of

(37)
$$
A(t,\rho) := \frac{1}{\sqrt{\coth^2 t - \rho}} \log \frac{\sqrt{\coth^2 t - \rho} + \sqrt{1 - \rho}}{\sqrt{\coth^2 t - \rho} - \sqrt{1 - \rho}}
$$

as $k \to \infty$, where t may depend on k. As will be seen below, the factors of as $\kappa \to \infty$, where ι may depend on κ . As will be seen below, the factors of $(\sqrt{2\pi t})^{-1}e^{-t/2}$ in front of the integral are always cancelled, leaving the prefactor $\frac{\sqrt{1-\rho}}{\rho}$. By (43), for pairs $(z, z + \frac{u}{\sqrt{k}})$, we have $\rho \to e^{-|u|^2}$, and $A_k(t, \rho)$ has a limit, which depends on whether or not we also send $t \to \infty$.

4.3. The limit as $k \to \infty$ for fixed *t*

This regime corresponds to letting the Brownian motion with respect to the Cartan-Killing metric evolve for a time t , and as discussed in $\S 3.3$ the ball of radius t in $g_{CK,k}$ metric is shrinking in size with respect to the Mabuchi metric and has d_k -radius equal to $\varepsilon_k t$.

The asymptotics in this regime could in principle be derived from the formula (8) of [9]. But more explicitly, we note that $\rho \to 0$ as $k \to \infty$ off the diagonal. Expanding at small ρ , we get

(38)
$$
\frac{1}{\sqrt{2\pi t}}e^{-t/2} \int_{-\infty}^{\infty} d\lambda \, e^{-\frac{1}{2t}\lambda^2} \frac{\cosh \lambda}{\sqrt{\coth^2 \lambda - \rho}}
$$

$$
\log \frac{\sqrt{\coth^2 \lambda - \rho} + \sqrt{1 - \rho}}{\sqrt{\coth^2 \lambda - \rho} - \sqrt{1 - \rho}} = 2t + \mathcal{O}(\rho),
$$

and the first term here cancels the first term in (36). Thus in the regime when we hold (z_1, z_2) fixed then $\rho \to 0$, we get

$$
I_{2,k}(t,\rho)\simeq a_0(t)+a_1(t)\rho+\cdots,
$$

where a_0, a_1, \ldots are constants independent of ρ . To obtain the 2-point correlation function of the Kähler metric, we then take four more derivatives. The constant a_0 does not contribute to the answer and we see that the two point correlation function is the free background term $\omega_{\phi_I}(z_1)\omega_{\phi_I}(z_2)$ plus a term exponentially decaying off the diagonal like $C_2(t)e^{-kD_I(z_1,z_2)}$.

4.4. The limit as $t \to \infty$ for fixed k

Now we apply steepest descent to the second integral as $t \to \infty$, and keeping k, N_k fixed. We obtain

$$
\partial_{\rho}I_{2,k}(\infty,\rho) := \lim_{t \to \infty} \partial_{\rho}I_{2,k}(t,\rho) = \lim_{t \to \infty} \frac{2t}{\rho} - \frac{1}{\rho} \left(2t + \log(1-\rho) + \mathcal{O}(1/t)\right)
$$

$$
= -\frac{\log(1-\rho)}{\rho}.
$$

Thus we have,

(39)
$$
I_{2,k}(\infty,\rho) = \text{Li}_2(\rho).
$$

As mentioned above, this is the same formula as (12). In the scaling limit around the diagonal with pairs of points of the form $(z, z + \frac{u}{\sqrt{k}})$ we have $\rho \simeq e^{-|u|^2}$. In the next section we connect this limit with the correlations between zeros of Gaussian random holomorphic sections, see (55) for a more precise statement.

4.5. The metric scaling limit with $t \to t \varepsilon_k^{-2}$

The goal now is to evaluate $I_{2,k}(\varepsilon_k^{-2}t,\rho)$ asympotically as $k \to \infty$. This scal-
ing league the d halls of uniform size as $k \to \infty$ with respect to the limit ing keeps the d_k -balls of uniform size as $k \to \infty$ with respect to the limit Mabuchi metric. Thus, as k changes the Brownian motion with respect to g_k probes distances of size t from the initial metric ω_0 for all k.

In order to apply steepest descent for $\lambda \in [0, \infty]$ we change variables $\lambda \to \varepsilon_k^{-2} \lambda$ in (36) so that the exponent becomes

$$
\varepsilon_k^{-2}\left(-\frac{\lambda^2}{2t}+\lambda\right).
$$

The saddle point occurs at $\lambda = t$, and the critical value of the exponent is $\frac{t}{2}$. The prefactor $e^{-t/2}$ in front of the integral, with $t \to \varepsilon_k^2 t$ again cancels the critical value of the phase and the singular term. The integrand in (36) at the saddle point is asymptotic to

(40)
\n
$$
\frac{1}{\sqrt{\coth^2(\varepsilon_k^{-2}t) - \rho}} \log \frac{\sqrt{\coth^2(\varepsilon_k^{-2}t) - \rho} + \sqrt{1 - \rho}}{\sqrt{\coth^2(\varepsilon_k^{-2}t) - \rho} - \sqrt{1 - \rho}}
$$
\n
$$
\simeq \frac{2\varepsilon_k^{-2}t}{\sqrt{1 - \rho}} + \frac{1}{\sqrt{1 - \rho}} \log(1 - \rho) + \cdots,
$$

so

$$
\partial_{\rho} I_{2,k}(\varepsilon_k^{-2}t,\rho) \simeq -\frac{\log(1-\rho)}{\rho},
$$

exactly as in (39).

4.6. Spatial scaling

In this section, we consider the scaling asymptotics of $I_{2,k}(\rho)$ discussed in §1.1. The only new element in the calculation is the scaling asymptotics of the Berezin kernel ρ . The calculation of $I_{2,k}(\rho)$ above does not change.

The main input into the scaling asymptotics is the following facts about the Berezin kernel (see [17, 18] for background and references). Off the diagonal one has

$$
\rho \simeq e^{-kD_I(z_1, z_2)},
$$

where the Calabi diastasis function is given by

(42)
$$
D_I(z_1, z_2) = \phi_I(z_1, z_1) + \phi_I(z_2, z_2) - \phi_I(z_1, z_2) - \phi_I(z_2, z_1).
$$

Here $\phi_I(z, \bar{z})$ is the Kähler potential (17) of the background Bergman metric ω_{ϕ_I} , corresponding to the identity matrix $P = I$, and $\phi_I(z_1, z_2)$ is its offdiagonal analytic extension.

The Berezin kernel has a scaling limit on the 'near-diagonal' where the distance $d(z_1, z_2)$ between z_1 and z_2 satisfies an upper bound $d(z_1, z_2) \leq$ $b\left(\frac{\log k}{k}\right)^{1/2}$ $(b \in \mathbb{R}^+)$. By comparison, $\rho \to 0$ rapidly on the 'far off-diagonal' where $d(z_1, z_2) \ge b \left(\frac{\log k}{k}\right)^{1/2}$, in the sense that for all $b, R > 0$,

$$
\nabla^j P_k(z_1, z_2) = O(k^{-R}) \quad \text{ uniformly for } d(z_1, z_2) \ge b \sqrt{\frac{\log k}{k}}.
$$

Here, ∇^j stands for the *j*-th covariant derivative.

The scaling asymptotics of this kernel near the diagonal may be described as follows. Let $z \in M$. Then

(43)
$$
P_k\left(z+\frac{u}{\sqrt{k}},z+\frac{v}{\sqrt{k}}\right)=e^{-\frac{1}{2}|u-v|^2}[1+R_k(u,v)],
$$

where

$$
|R_k(u, v)| \le \frac{C_2}{2} |u - v|^2 k^{-1/2 + \varepsilon}, \quad |\nabla R_k(u)| \le C_2 |u - v| k^{-1/2 + \varepsilon},
$$

$$
|\nabla^j R_k(u, v)| \le C_j k^{-1/2 + \varepsilon} \quad j \ge 2,
$$

for $|u| + |v| < b\sqrt{\log k}$.

It follows that the scaling asymptotics of the variance term for random metrics is given by

(44)
$$
I_{2,k}(\varepsilon_k^{-2}t, P_k(z, z + u/\sqrt{k})) \simeq \text{Li}_2(e^{-|u|^2}),
$$

just as in the limit as $t \to \infty$ first, considered in §4.4.

4.7. Energy entropy scaling

Another natural scaling comes directly from the density (24). We separate out the 'action' from the 'amplitude' and express it in terms of the empirical measures $d\mu_{\lambda} = \frac{1}{N_k} \sum_{j=1}^{N_k} \delta_{\lambda_j}$ of the eigenvalues of Λ

(45)
$$
\Delta(\lambda)\Delta(e^{\lambda})e^{-\frac{1}{4t}\sum_{j=1}^{N}\lambda_j^2}
$$

= $e^{N_k^2(\int_M \int_M \log|x-y|d\mu_{\lambda}(x)d\mu_{\lambda}(y)+(\int_M \int_M \log|e^x-e^y|d\mu_{\lambda}(x)d\mu_{\lambda}(y)))+\frac{N_k}{2t}\int_M x^2d\mu_{\lambda}}.$

To give all terms the same order in N_k we need to rescale the time to $t \to t/N_k$. Then the scaled measures μ_k satisfy a large deviations principle with the rate function given by the exponent of (45).

5. The geodesic boundary of \mathcal{B}_k and configurations of zeros

In this section, we explain how the result of §4.4 is essentially the same as the theory of zeros of random holomorphic sections. As $t \to \infty$, the mass of the heat kernel gets concentrated on a part of the boundary of \mathcal{P}_{N_k} corresponding to 'singular metrics' given by zero sets of holomorphic sections.

Symmetric spaces of non-compact type have several different notions of boundary and several types of compactifications. For background we refer to [11]. The boundary relevant to the heat kernel measures and their $t \to \infty$ limit is best stated in terms of the Bergman metrics themselves and their limits along geodesic rays of \mathcal{P}_{N_k} .

Definition 1. The weak* compactification of \mathcal{B}_k is $\mathcal{B}_k \cup \partial \mathcal{B}_k$ where $\partial \mathcal{B}_k$ is the set of limit points (i.e. endpoints) of the Bergman metrics along Bergman geodesic rays $\omega_k(s)$.

In fact, the only relevant boundary points are the ones arising from the geodesic ray starting at the balanced metric $x_k \simeq 0$ and with initial velocity in the direction of δ_N , together with the endpoints of the $U(N_k)$ -orbit of this ray.

A geodesic ray in the space of Bergman potentials is a one-parameter family of metrics whose potentials have the form,

$$
\beta_t = \frac{1}{k} \log \sum_j e^{t\lambda_j} |s_j^U(z)|^2,
$$

where in $SL(N_k, \mathbb{C})/SU(N_k)$ the ray starts at the origin and has initial vector (U, Λ) . We note that

(46)
$$
\frac{1}{k} \log \sum_{j} e^{t\lambda_j} |s_j^U(z)|^2
$$

$$
= \frac{t\lambda_{\max}}{k} + \frac{1}{k} \log \left(|s_{\max}^U(z)|^2 + \sum_{j \neq \max} e^{t(\lambda_j - \lambda_{\max})} |s_j^U(z)|^2 \right).
$$

Here, λ_{max} is the largest of the λ_j and s_{max}^U is the corresponding section. Clearly,

(47)
$$
\frac{1}{k} \log \sum_{j} e^{t\lambda_j} |s_j^U(z)|^2 - \sup_{M} \frac{1}{k} \log \sum_{j} e^{t\lambda_j} |s_j^U(z)|^2
$$

$$
= \frac{1}{k} \log \left(|s_{\max}^U(z)|^2 + \sum_{\lambda_j \neq \max} e^{t(\lambda_j - \lambda_{\max})} |s_j^U(z)|^2 \right)
$$

$$
- \sup_{M} \frac{1}{k} \log \left(|s_{\max}^U(z)|^2 + \sum_{\lambda_j \neq \max} e^{t(\lambda_j - \lambda_{\max})} |s_j^U(z)|^2 \right).
$$

This family of potentials is bounded above by 0 and is pre-compact in $L^p(M)$ for all $1 \leq p \leq \infty$. The sup in the second term is itself bounded above by

 $\frac{\log k}{k}$ and therefore we may remove the sup without changing the limit (i.e. the sup was only needed to get rid of the $\frac{t\lambda_{\max}}{k}$ term). We then observe that for any Bergman geodesic ray,

(48)
$$
\lim_{t \to \infty} \left\| \frac{1}{k} \log \left(|s_{\max}^U(z)|^2 + \sum_{\lambda_j \neq \max} e^{t(\lambda_j - \lambda_{\max})} |s_j^U(z)|^2 \right) - \frac{1}{k} \log |s_{\max}(z)|^2 \right\|_{L^1(M)} \to 0.
$$

Indeed,

(49)
$$
F_t(z) := \left(|s_{\max}^U(z)|^2 + \sum_{\lambda_j \neq \max} e^{t(\lambda_j - \lambda_{\max})} |s_j^U(z)|^2 \right)
$$

is monotonically decreasing to $|s_{\text{max}}^{U}(z)|^2$ for each z. Therefore its logarithm
monotonically decreases to log $|s_{\text{max}}^{U}(z)|^2$. If we subtract $\frac{1}{z} \log |s_{\text{max}}(z)|^2$ than monotonically decreases to $\log |s_{\text{max}}^{U}(z)|^{2}$. If we subtract $\frac{1}{k} \log |s_{\text{max}}(z)|^{2}$ then
the difference is always ≥ 0 and we may remove the absolute values, and the difference is always ≥ 0 and we may remove the absolute values, and then apply the monotone convergence theorem to take the limit under the integration sign. But the limit equals zero almost everywhere since each term tends to zero, proving (48).

We now consider the $(1, 0)$ forms obtained by taking ∂ of the potentials. Taking ∂ kills the $\frac{t\lambda_{\max}}{k}$ term, and gives,

(50)
$$
\frac{1}{k} \partial \log \sum_{j} e^{t\lambda_j} |s_j^U(z)|^2
$$

$$
= \frac{1}{k} \partial \log \left(|s_{\max}^U(z)|^2 + \sum_{\lambda_j \neq \max} e^{t(\lambda_j - \lambda_{\max})} |s_j^U(z)|^2 \right).
$$

Proposition 5.1. The weak limits of Bergman metrics along geodesic rays are generically given by the normalized zero distributions of holomorphic sections of L^k , i.e. as $t \to \infty$,

$$
\frac{1}{k}\partial\bar{\partial}\log\sum_{j}e^{t\lambda_j}|s_j^U(z)|^2\to \frac{1}{k}\partial\bar{\partial}\log|s|^2.
$$

If the highest weight has multiplicity r, then the limit is $\frac{1}{k} \partial \bar{\partial} \log \sum_{j=1}^{r} |s_j|^2$
where $1 \leq r \leq r$ and $\lceil s_1 \rceil^r$ is any set of sections in $H^0(M, I^k)$ where $1 \leq r \leq n$ and $\{s_j\}_{j=1}^r$ is any set of sections in $H^0(M, L^k)$.

Proof. We take $\partial\bar{\partial}$ of the two terms of (48) to obtain,

(51)
$$
\frac{1}{k}\partial\bar{\partial}\log\sum_{j}e^{t\lambda_{j}}|s_{j}^{U}(z)|^{2} \to \frac{1}{k}\partial\bar{\partial}\log|s_{\max}^{U}(z)|^{2}
$$

in the sense of distributions. That is, if we integrate against a smooth $(n 1, n - 1$) form ψ (i.e. a smooth function if the dimension of M equals one), then

$$
\int_M \psi \wedge \frac{1}{k} \partial \bar{\partial} \log \left(|s_{\max}^U(z)|^2 + \sum_{j \neq \max} e^{t(\lambda_j - \lambda_{\max})} |s_j^U(z)|^2 \right)
$$

$$
= \int_M \partial \bar{\partial} \psi \wedge \frac{1}{k} \log \left(|s_{\max}^U(z)|^2 + \sum_{\lambda_j \neq \max} e^{t(\lambda_j - \lambda_{\max})} |s_j^U(z)|^2 \right)
$$

and by (48) the right side tends $\int_M \partial \overline{\partial} \psi \wedge \frac{1}{k} \log |s_{\max}^U(z)|^2$.
Multiplicity *x* moops that λ has multiplicity *x* and

Multiplicity r means that λ_{max} has multiplicity r, and then one sums r the associated sections s. over the associated sections s_i .

A key point for this article is that each of the weights in δ_N is distinct. Hence $r = 1$ in that direction, and therefore the boundary points in the direction δ_N and its $U(N)$ -translates all consist of δ -functions $\frac{1}{k} \partial \bar{\partial} \log |s(z)|^2$ of holomorphic sections of L^k . Every section arises in the $U(N_k)$ -orbit.

5.1. Relation to zeros of holomorphic sections

In [17], the authors found a bi-potential $Q_N \in L^1(M \times M)$ for the pair correlation function of zeros of Gaussian random holomorphic sections

(52)
$$
d\gamma(s) = \frac{1}{\pi^m} e^{-|c|^2} dc, \quad s = \sum_{j=1}^n c_j s_j,
$$

on L^k , where $\{s_i\}$ is an orthonormal basis and dc is 2n-dimensional Lebesgue measure. This Gaussian is characterized by the property that the $2n$ real variables Re c_j , Im c_j $(j = 1, \ldots, n)$ are independent Gaussian random variables with mean 0 and variance $\frac{1}{2}$; i.e.,

$$
\mathbf{E}c_j = 0, \quad \mathbf{E}c_jc_k = 0, \quad \mathbf{E}c_j\bar{c}_k = \delta_{jk}.
$$

The current of integration Z_s over the zeros of one section is given by the Poincaré-Lelong formula

(53)
$$
Z_s = \frac{\sqrt{-1}}{\pi} \partial \bar{\partial} \log |f| = \frac{\sqrt{-1}}{\pi} \partial \bar{\partial} \log ||s^N||_h + \phi_I(z) ,
$$

As mentioned above, the limit as $t \to \infty$ of random Bergman metrics along the rays above must be random singular metrics, and we claim that the limit ensemble is equivalent to the Gaussian one. Indeed, the limit measure is $U(N_k)$ -invariant and there exists just one such measure up to equivalence, namely the Gaussian measure above. This explains why the bipotential $4\pi^2 Q_k$ of [17] is the same as the $t \to \infty$ limit of the two-point function $I_{2,k}(\rho)$ of the heat kernel ensemble in §4.4, corrobrating that this limit ensemble is that of random zeros of sections.

5.2. Smooth linear statistics and asymptotic normality

Similar to the area random variable X_U but easier to work with, is the smooth linear statistic

(54)
$$
X_f(\omega) = \int_M f\omega,
$$

where $f \in C^{\infty}(M)$ is a smooth test function. It has a much smaller variance than X_U since the effect of 'zeros along the boundary' is smoothed out. In [18] it is proved that

(55)
$$
Var(X_f) = k^{-1} \left[\frac{\zeta(3)}{16\pi} ||\Delta f||_2^2 + O(k^{-\frac{1}{2} + \varepsilon}) \right].
$$

It was proved by Sodin-Tsirelson [19] in certain cases and then [17] for the present setting that the random variables X_f^k obey an asymptotic central limit theorem. Namely, if we normalize X_f to have mean zero and variance one, then

(56)
$$
\frac{X_f - \mathbf{E}(X_f)}{\sqrt{\text{Var}(X_f)}}
$$

tends in distribution to the standard Gaussian distribution $\mathcal{N}(0,1)$ as $k \to$ ∞ . That is,

$$
k^{1/2}\left(X_f - \frac{k}{\pi} \int_M f\omega_\phi\right) \to \mathcal{N}(0, \sqrt{\kappa_1} \|\partial \bar{\partial} f\|_2)
$$

in the weak sense of convergence of distributions as $k \to \infty$, where κ_1 is a certain positive universal constant and $\mathcal{N}(0, \sigma)$ denotes the (real) Gaussian distribution of mean zero and variance σ^2 .

The same results hold in the limiting case of the heat kernel measures when $t \to \infty$. It would be interesting to investigate the analogous variance and asymptotic normality results for heat kernel measures in the other regimes as $k \to \infty$, $t \to \infty$. The formulae of this article combined with the techniques of [17, 18] make this possible.

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MATHEMATISCHES INSTITUT, UNIVERSITÄT ZU KÖLN WEYERTAL 86-90, 50931 KÖLN, GERMANY & Simons Center for Geometry and Physics, Stony Brook University NY 11794, USA $\it E\mbox{-}mail\;address:$ sam.klevtsov@gmail.com

Department of Mathematics, Northwestern University Evanston, IL 60208, USA & Simons Center for Geometry and Physics, Stony Brook University NY 11794, USA E-mail address: zelditch@math.northwestern.edu